# Workshop on "Science with High-Power Lasers and Pulsed Power (#3) Research Opportunities and User Meeting"





Santa Fe Eldorado Hotel July 28-30, 2011 Presented by:

Alan Wootton
University of Texas

#### Introduction

- Purpose, deliverables of this meeting
- •Summary of 2009 (inaugural,) 2010 workshops and consequences
  - Status of highlighted concepts
- Agenda
  - -User meeting:
    - -Ongoing 'user' research at the Z facilities and TX lasers
    - -The facilities and their capabilities
  - -Research Directions
    - -Breakout sessions. Discuss
      - -How to maximize the impact of our current research
      - -How to grow the community and where appropriate the teams
      - -New ideas
      - -Infrastructure requests
  - -Discussion
- Administrative information

# The primary purpose of this meeting is to help maximize the impact of our fundamental science program

A workshop to discuss fundamental science experiments that are being or about to be, or could be, undertaken on the pulsed power and laser facilities at Sandia National Laboratories (SNL) and the University of Texas (UT).

A user meeting, in which current and prospective users can present their research, and hear plans for and comment on user access, planned experiments, facility capabilities, and facility schedules.

Up to 15% of the shot time on the Z facilities is now available for fundamental science. The Texas Petawatt laser is operating and accepting 'collaborative users'.

The workshop is being held under the auspices of the Institute for High Energy Density Science, joint between UTX and SNL.

Maximize the impact of our research, while being realistic in what can be achieved with limited shot time on these unique facilities ('manage, maintain expectations')

#### The facilities considered are at SNL and UTX

At SNL: the Z accelerator with its experimental platforms, and diagnostics (including Z-Beamlet and Z-Petawatt) are available for HED science experiments. Up to 15% of Z shots will be available for fundamental science. The program is now run by Mark Herrmann of SNL.

At UTX (Austin): Texas Petawatt, THOR, GHOST. Mike Donovan is point of contact for the collaborative user program.

Information on diagnostics, facility capabilities, and plasma platforms, will be provided this afternoon.

### There are five deliverables expected from this workshop

- 1) A status report for each research project with allocated experimental time (astrophysics, planetary science, magnetized HED science)
- 2) A summary of how each individual research proposal will maximize the impact of the research while managing expectations (limited shots, diagnostics, etc.)
- 3) Specific ideas on how to grow our community and where appropriate the individual research teams

-----

- 4) New research directions for consideration. However we need our committed-to fundamental research to succeed, so this must have priority
- 5) Infrastructure suggestions (diagnostics, design tools, analysis tools, data access, lab space, ....). However initial focus should be on utilizing existing infrastructure

Each breakout session should prepare an informal report for Saturday morning covering the above, and a more formal report for our web page

# In two years we have progressed from first ideas to initiating experiments and preliminary data



Science with High-Power Lasers and Pulsed Power (#2), Research Opportunities and User Meeting





**2009:** The first workshop considered research directions, defined four research areas (astro, planetary, magnetized HED, particles and beams) with 10 ideas.

**2010:** The second workshop was both a research directions and a user meeting. Within the four areas, eight specific research proposals were initiated. Additional diagnostics, capabilities and infrastructure were suggested.

Proposals were submitted to the HEDLP Joint program.

A draft document planning for users was developed.

At SNL: A call for proposals to utilize the Z facilities for fundamental science was published.

**2011:** At SNL, received proposals were reviewed, and time has been allocated.

At UTX: Collaborative user experiments allocated time on the TX PW laser.

We plan to continue the research directions and user meeting, and calls for proposals, with staggered project starts to maintain continuity

### Agenda, Thursday July 28

- <u>Session 1a: User experiments: status at Z</u>. Chaired by Alan Wootton (UTX)
  - White dwarf work status: Don Winget (UTX)
  - Opacity work: Anil Pradham (Ohio St)
  - Solar opacity work status: Jim Bailey (SNL)
  - Photo-ionized plasma work status: Roberto Mancini (UNR)
  - Black hole accretion disk status: Duane Liedahl (SNL)
  - Giant Planets: far out, close in, and deep inside: Nadine Nettelmann (U. Rostock)
  - Earth core work status: Stein Jacobsen (Harvard)
- Session 1b: User experiments: status at UT. Chaired by Mike Donovan (UTX)
  - Solid Target Experiments on the Texas Laser Facilities: Kramer Akli and Mike Storm (Ohio St)
  - Cluster fusion work status: Gilliss Dyer (UTX)
  - Wakefield acceleration work status: Xiaoming Wang (UTX)
  - Antimatter work status: Edison Liang (Rice)
- Session 2: Facility access, facility status. Chaired by Briggs Atherton (SNL)
  - Z diagnostics and plasma platforms: Greg Rochau (SNL)
  - UT user capabilities: Mike Donovan (UTX)

### Agenda, Friday July 29 (breakout groups) and Saturday July 30

#### Friday July 29:

- 3.1) Magnetized HED Science, chaired by Dan Sinars (SNL,) in Pinon Room (2<sup>nd</sup> floor)
- 3.2) Planetary Science, chaired by Thomas Mattsson (SNL), in Chaparral Room (3<sup>rd</sup> floor)
- 3.3) Astrophysics, chaired by Don Winget (UTX) and Jim Bailey (SNL), in Sunset Room (here)

Spare room for any other group that self-organizes: Turquoise Room (4th floor)

#### Sat July 30: Sunset room (here)

Out-brief from panel chairs (Dan Sinars, Thomas Mattsson, Jim Bailey)
General discussion

#### **Administrative Details**

Meeting Room Assignments (also listed on your printed materials):

Plenary Sessions: Sunset

Parallel Breakouts Sessions (Friday all day):

Astrophysics: Sunset (here)

Planetary: Chaparral Room, third floor

Magnetized HED: Pinon room, second floor

Spare: Turquoise, fourth floor

(You cannot work in the break-out rooms past 6 pm)

Note start times: Thursday & Friday breakfast 7:30 – 8:30, start time 8:30; Saturday breakfast 8:00 – 9:00 am, start time 9:00 am

Schedule changes:

Today: Please hand in presentations to me at the break before your presentation We will ask for cc of the presentations, to post on the web page

Maria: maria.aguirre@mail.utexas.edu

Cari: cagerlo@sandia.gov (505) 417-2398

# Back up slides

# Research areas and topics from workshops #1 and #2

Science area	research topic	Ongoing research	New idea	Community	
Radiative astrophysics	White dwarf photoshperes	Lab (Z) spectral properties in H	Study (in Z) He, C. Add B fields. Lower temperatures	Add: postdoc at Z, spectroscopis t, line theorist	
	Accretion- powered objects	Test spectral models used to infer driving radiation and absorption in active galactic nuclei	Test and improve spectral synthesis models associated with accretion disks around black holes	Need PI. Add: postdoc, student at Z; improve links to astro community	
	Stellar opacities	Fe opacity in sun's convection zone	Mid-Z opacities	Add: postdoc a Z; additional theoretical effort	
Particles and beams	Proton beams	Ddeflectometry diagnostic development	Distinguish roles of different electron acceleration models, including multi-pass through target	Add: more theoretical effort	
	Electron beams	Wakefield acceleration			
Magnetized high energy density science	Hydrodynamic, magnetized jets	At MAGPIE, GenASIS, x- pinch	Z -> highter dimensionless numbers (Reynolds), resolution, density (all for turbulence studies). X-pinches-> B -> pulsars?	Consider -possibility of LANL -National High -Field Lab, others	
	Boundary effects	x-pinch	Higher current at Z; two-fluid modeling		
	Cluster fusion Kelvin- Helmholtz instability Reconnection	no B field	With 100 T (energy transport)  In non-ideal warm dense matter		
Planetray science	Earths and super-earths		-> better diagnostics EOS up to 400 GPa and 5000 K in Fe and alloys with impurities- > a) Earth now, understanding dynamo generation, b) Earth then, understanding moon formation	Grow teams, embrace geographic diversity	
	gas and ice giants here and out there	at breakout sessions	High-precision EOS and density for H, H2O above 100 Gpa and 5000 K. Study mixtures and phase stability		

# Diagnostic and capability suggestions from workshops #1 and #2

	radiative astrophysics	particles and beams	magnetized high energy density	planetary science			
DIAGNOSTICS							
optical Thomson scattering (n, T)	х		х				
x-ray Thomson scattering (n, T)	×		х	х			
laser interferometer (n)	Х	X					
high sensitivity spectroscopy (emission spectra)	×						
proton beam deflectometry (B)		х	х				
pyrometry				Х			
proton radiography (n)	X						
high intensity backlighters	X						
spectrometer 0.5-1 keV	X						
improved single-shot laser characterization		×					
laser Compton scattering		X					
Stark shift (B)		X					
CAPABILITIES							
filed gas cells on top of x-ray source (higher photoionization parameter)	x						
cells with hazardous gasses	X						
thinner windows on cells	Х						
longer duration x-ray drives (for equilibrium conditions)	х						
x-pinch geometry			Х				
Table 2. Diagnostic and other	capabilities req	uested					

## Infrastructure suggestions from workshops #1 and #2

- 1) Providing better infrastructure for visitors (offices, network, computing, 'phone, project meeting rooms/collaborative environment)
- 2) Developing a well-defined process for machine-time application
- 3) List of diagnostics: past, present and future
- 4) A Calculator" for simulating existing diagnostics and predicting experiment outcomes similar to what is available for space and ground based observatory instruments
- 5) A POC, but from a hardware perspective: someone who knows all the hardware, what is available and where it is and who to talk to